

Lecture 11-3

Radiation
from a moving
point charge

Liénard-Wiechert
potentials
Fields

Power radiated by a
point charge

Larmor formula

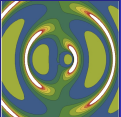
PHYSICS 453

Electromagnetism II

Lecture 11-3

Physics Department
Old Dominion University

March 26, 2026



Outline

Lecture 11-3

Radiation
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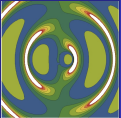
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- 1 Radiation from a moving point charge
 - Liénard-Wiechert potentials
 - Electric and Magnetic Fields
 - Power radiated by a point charge
 - Larmor formula



Liénard-Wiechert potentials

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- The Liénard-Wiechert potentials for a point charge are

$$\Phi(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{1}{|\mathbf{r} - \mathbf{w}(t_r)| - \mathbf{v}(t_r) \cdot (\mathbf{r} - \mathbf{w}(t_r))/c}$$

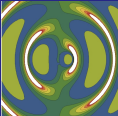
$$\mathbf{A}(\mathbf{r}, t) = \frac{\mathbf{v}(t_r)}{c^2} \Phi(\mathbf{r}, t)$$

- Introducing the notation $\boldsymbol{\varsigma}(t) \equiv \mathbf{r} - \mathbf{w}(t)$, we obtain

$$\Phi(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{1}{\boldsymbol{\varsigma}(t_r) - \mathbf{v}(t_r) \cdot \boldsymbol{\varsigma}(t_r)/c}$$

$$\mathbf{A}(\mathbf{r}, t) = \frac{\mathbf{v}(t_r)}{c^2} \Phi(\mathbf{r}, t) = \frac{q\mu_0}{4\pi} \frac{\mathbf{v}(t_r)}{\boldsymbol{\varsigma}(t_r) - \mathbf{v}(t_r) \cdot \boldsymbol{\varsigma}(t_r)/c}$$

- According to these formulas, the field at the point of observation at time t is determined by the state of motion of the charge at the earlier time t_r
- Also, $\boldsymbol{\varsigma}(t) = \mathbf{r} - \mathbf{w}(t)$ is the radius vector from the charge q to the observation point P ; like $\mathbf{w}(t)$ it is a given function of the time
- The time t_r is determined by the equation $t_r + \boldsymbol{\varsigma}(t_r)/c = t$



Electric and Magnetic Fields, cont.

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- In Lecture 10-3 we obtained the corresponding expressions for the fields

$$\mathbf{E} = \frac{q}{4\pi\epsilon_0 (\varsigma - \boldsymbol{\varsigma} \cdot \mathbf{v}/c)^3} \left\{ (1 - v^2/c^2) \left(\boldsymbol{\varsigma} - \frac{\mathbf{v}}{c} \varsigma \right) + \frac{1}{c^2} \boldsymbol{\varsigma} \times \left[\left(\boldsymbol{\varsigma} - \frac{\mathbf{v}}{c} \varsigma \right) \times \mathbf{a} \right] \right\}$$

$$\mathbf{B} = \frac{1}{\varsigma} \boldsymbol{\varsigma} \times \mathbf{E}$$

Here, $\mathbf{a} = \partial \mathbf{v} / \partial t_r$

- All quantities on the right sides of the equations refer to the time t_r
- Note that the magnetic field is everywhere perpendicular to the electric
- In the non-relativistic limit, the electric field $\mathbf{E}$ reduces to the Coulomb field

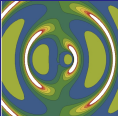
$$\mathbf{E}(\mathbf{r}, t)|_{v/c \rightarrow 0} \rightarrow \frac{q\boldsymbol{\varsigma}}{4\pi\epsilon_0\varsigma^3},$$

while the magnetic field tends to zero as v/c

- Introduce the unit vector $\hat{\boldsymbol{\varsigma}} \equiv \boldsymbol{\varsigma}/\varsigma$, and also the notation $\mathbf{u} \equiv \hat{\boldsymbol{\varsigma}} - \mathbf{v}(t_r)/c$
- This vector tends to $\hat{\boldsymbol{\varsigma}}$ in the non-relativistic limit $v/c \rightarrow 0$
- Using $\boldsymbol{\varsigma} - \boldsymbol{\varsigma} \cdot \mathbf{v}/c = (\boldsymbol{\varsigma} \cdot \mathbf{u})$, we can write the electric and magnetic fields as

$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{\boldsymbol{\varsigma}}{(\boldsymbol{\varsigma} \cdot \mathbf{u})^3} \left[\mathbf{u}(1 - v^2/c^2) + \boldsymbol{\varsigma} \times (\mathbf{u} \times \mathbf{a})/c^2 \right]$$

$$\mathbf{B}(\mathbf{r}, t) = \hat{\boldsymbol{\varsigma}} \times \mathbf{E}(\mathbf{r}, t)/c$$



Electric and Magnetic Fields, cont.

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$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{\zeta}{(\zeta \cdot \mathbf{u})^3} \left[\mathbf{u}(1 - v^2/c^2) + \zeta \times (\mathbf{u} \times \mathbf{a})/c^2 \right]$$

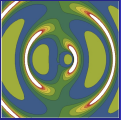
$$\mathbf{B}(\mathbf{r}, t) = \frac{\hat{\zeta}}{c} \times \mathbf{E}(\mathbf{r}, t)$$

- As usual, velocity and acceleration are taken at $t = t_r$.
(recall that t_r is defined as a solution to the equation $c(t - t_r) = \zeta$)
- The electric field consists of two parts of different type:
the first term ($\sim \mathbf{u}$) is called the **velocity field**, and
the second ($\sim \mathbf{a}$) is called the **acceleration or the radiation field**.
- The first term varies at large distances like $1/\zeta^2$.
Since it is independent of the acceleration, it corresponds to the field produced by a uniformly moving charge. For large distances ζ , the velocity field decreases as $1/\zeta^2$
- The total field $\mathbf{E}(\mathbf{r}, t)$ contains also a radiation component $\mathbf{E}^{\text{rad}}(\mathbf{r}, t)$. It decreases as $1/\zeta$ for large ζ , and is given by

$$\mathbf{E}_{\text{rad}}(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0 c^2 \zeta} \frac{\hat{\zeta} \times (\mathbf{u} \times \mathbf{a})}{(\hat{\zeta} \cdot \mathbf{u})^3} = \frac{q}{4\pi\epsilon_0 c^2 \zeta} \frac{\hat{\zeta} \times (\mathbf{u} \times \mathbf{a})}{(1 - \hat{\zeta} \cdot \mathbf{v}/c)^3}$$

- The radiation magnetic field also decreases as $1/\zeta$ for large ζ

$$\mathbf{B}_{\text{rad}}(\mathbf{r}, t) = \frac{\hat{\zeta}}{c} \times \mathbf{E}_{\text{rad}}(\mathbf{r}, t)$$



Power radiated by a point charge

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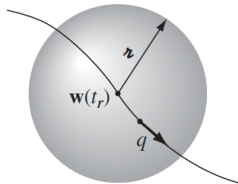
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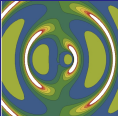
- The Poynting vector for the radiation component is

$$\mathbf{S} = \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B} = \frac{1}{\mu_0 c} \mathbf{E} \times (\hat{\boldsymbol{\zeta}} \times \mathbf{E}) = \frac{1}{\mu_0 c} [E^2 \hat{\boldsymbol{\zeta}} - (\hat{\boldsymbol{\zeta}} \cdot \mathbf{E}) \mathbf{E}]$$

- Some of the energy is radiation; another part is just a field energy carried along by the particle as it moves
- To calculate the power radiated by the particle at time t_* , we draw a large sphere with radius $\varsigma = R$, wait for $t - t_* = \frac{R}{c}$, and integrate Poynting vector over the surface



- The velocity field is $\sim 1/R^2$. The corresponding P_{rad} is $\sim R^2 \frac{1}{R^4} = \frac{1}{R^2}$ so it does not contribute to the radiated power at large R



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$$\mathbf{S} = \frac{1}{\mu_0 c} [E^2 \hat{\boldsymbol{\zeta}} - (\hat{\boldsymbol{\zeta}} \cdot \mathbf{E}) \mathbf{E}]$$

- The power due to the acceleration field ($\sim 1/R$) is finite: $P_{\text{rad}} \sim R^2 \frac{1}{R^2} = 1$

$$\mathbf{E}_{\text{rad}}(\mathbf{r}, t) = \frac{q}{4\pi\epsilon_0 c^2 \zeta} \frac{\hat{\boldsymbol{\zeta}} \times (\mathbf{u} \times \mathbf{a})}{(\hat{\boldsymbol{\zeta}} \cdot \mathbf{u})^3}$$

$$\rightarrow \boldsymbol{\zeta} \cdot \mathbf{E}_{\text{rad}}(\mathbf{r}, t) = 0 \Rightarrow \mathbf{S}_{\text{rad}} = \frac{\hat{\boldsymbol{\zeta}}}{\mu_0 c} E_{\text{rad}}^2$$

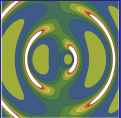
- For simplicity, consider the charge which is instantaneously at rest at $t = t_*$
- Since $\mathbf{v}(t_*) = 0$, $\mathbf{u}(t_*) = \hat{\boldsymbol{\zeta}}$, so

$$\mathbf{S}_{\text{rad}} = \frac{\hat{\boldsymbol{\zeta}}}{\mu_0 c} \left(\frac{\mu_0 q}{4\pi R} \right)^2 [a^2 - (\hat{\boldsymbol{\zeta}} \cdot \mathbf{a})^2] = \frac{\mu_0 q^2 a^2}{16\pi^2 c} \frac{\sin^2 \theta}{R^2} \hat{\boldsymbol{\zeta}}$$

- The total power is given by the following Larmor formula

$$P_{\text{rad}} = \oint_S \mathbf{S}_{\text{rad}} \cdot d\mathbf{S} = \frac{\mu_0 q^2 a^2}{16\pi^2 c R^2} \int \frac{\sin^2 \theta}{R^2} R^2 \sin \theta d\theta d\varphi = \frac{\mu_0 q^2 a^2}{6\pi c}$$

- We have already obtained this using the electric dipole radiation formulas



Larmor formula, cont.

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- We have derived the Larmor formula under the assumption that $v = 0$
- One can demonstrate that it holds true as long as $v \ll c$
- In the general case of arbitrary velocity, the radiation is given by the Lienard formula

$$P_{\text{rad}} = \frac{\mu_0 q^2 \gamma^6}{6\pi c} \left(a^2 - \frac{(\mathbf{v} \cdot \mathbf{a})^2}{c^2} \right)$$

where $\gamma \equiv 1/\sqrt{1 - \frac{v^2}{c^2}}$