

807 Midterm (20 points). 10/25/16, 10:50 a.m. - 12:20 p.m.

**Problem 1.**

A particle of mass  $m_1$  moving from infinity with velocity  $v_0$  and impact parameter  $b$  scatters off a target particle of mass  $m_2$  initially at rest. Find the minimum distance between the particles for the repulsive interaction potential

$$U(r) = \frac{\alpha}{r^2}$$

**Solution**

The problem is equivalent to the scattering of a particle with effective mass  $m = \frac{m_1 m_2}{m_1 + m_2}$  in the central field with  $U(r) = \frac{\alpha}{r^2}$ .

The kinetic energy of the “effective particle” is  $\frac{\mu v^2}{2}$  and the effective potential is

$$U_{\text{eff}} = \frac{\alpha}{r^2} + \frac{L^2}{2\mu r^2} =$$

where  $L = \mu v b$  is the angular momentum. At minimal distance  $r_{\min}$  there is no kinetic energy so the conservation of energy gives

$$\frac{\mu v^2}{2} = \frac{\alpha}{r_{\min}^2} + \frac{L^2}{2\mu r_{\min}^2} \Rightarrow r_{\min} = b \sqrt{1 + \frac{2\alpha}{\mu v^2 b^2}}$$

**Solution #2**

Let us solve this problem in c.m. frame  $m_1 \vec{r}_1 + m_2 \vec{r}_2 = 0$ ,  $\vec{r} = \vec{r}_1 - \vec{r}_2$ . We have a scattering of two particles: particle “1” at impact parameter  $b_1 = \frac{m_2}{M} b$  with velocity at  $\infty$   $v_{10} = \frac{m_2}{M} v_0$  and particle “2” with  $b_2 = -\frac{m_1}{M} b$  and  $v_{20} = -\frac{m_1}{M} v_0$

$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2} = \frac{m_1 \vec{r}_1}{m_1 + m_2} \quad \dot{\vec{R}} = \frac{m_1 \vec{v}}{m_1 + m_2}$$

where  $M \equiv m_1 + m_2$ . Conservation of energy reads

$$E = \frac{m_1 v_1^2}{2} + \frac{m_2 v_2^2}{2} + V(r)$$

Let us express it in terms of  $\mu$  and  $\vec{r}$  (and  $\vec{v} = \dot{\vec{r}}$ ). Since

$$\begin{aligned} \vec{r}_1 &= \frac{m_2}{M} \vec{r}, & \vec{v}_1 &= \frac{m_2}{M} \vec{v}, & \vec{r}_2 &= -\frac{m_1}{M} \vec{r}, & v_2 &= -\frac{m_1}{M} v, \\ L_1 &= m_1 v_{10} b_1 = \frac{m_2}{M} \mu v b, & L_2 &= m_2 v_{20} b_2 = \frac{m_1}{M} \mu v b \end{aligned}$$

we get

$$E = \frac{m_1 v_1^2}{2} + \frac{m_2 v_2^2}{2} = \frac{\mu v^2}{2} + V(r) = \frac{\mu \dot{r}^2}{2} + \frac{\mu r^2 \dot{\phi}^2}{2} + V(r)$$

Conservation of angular momentum reads

$$L_1 + L_2 = m_1 r_1^2 \dot{\phi} + m_2 r_2^2 \dot{\phi} = \mu r^2 \dot{\phi} = \mu v b \Rightarrow \dot{\phi} = \frac{vb}{r^2}$$

so we get

$$E = \frac{\mu \dot{r}^2}{2} + \frac{\mu r^2 \dot{\phi}^2}{2} + V(r) = \frac{\mu \dot{r}^2}{2} + \frac{\mu v^2 b^2}{2r^2} + \frac{\alpha}{r^2}$$

At  $r_{\min}$  we have  $\dot{r} = 0$  which again results in

$$\frac{\mu v^2}{2} = \frac{\alpha}{r_{\min}^2} + \frac{\mu v^2 b^2}{2r_{\min}^2} \Rightarrow r_{\min} = b \sqrt{1 + \frac{2\alpha}{\mu v^2 b^2}}$$

### Problem 2.

Imagine that you are on a large rotating disk, with the angular velocity  $\vec{\omega}$  perpendicular to the disk. Assume that you know that the center of the disk is fixed, no real forces act on anything on the disk, and  $\vec{\omega}$  changes only in magnitude. Suppose that at time  $t = 0$  you release a point mass  $m$  and observe that it moves away in the radial direction. If the angular velocity at  $t = 0$  was  $\omega_0$ , what is the magnitude and direction of Coriolis force at a later time  $t$ ?

### Solution # 1

If the point mass is moving radially the sum of forces in  $\hat{\phi}$  direction is zero:

$$2m(\vec{\omega} \times \dot{\vec{r}}) + m(\dot{\vec{\omega}} \times \vec{r})$$

Since  $\vec{\omega} \perp \vec{r}$

$$2\omega \dot{r} = -\dot{\omega} r \Rightarrow \frac{d}{dt} \omega r^2 = 0 \Rightarrow \omega r^2 = \omega_0 a^2 \Rightarrow r(t) = a \sqrt{1 + \omega_0^2 t^2}$$

so Newton's 2nd law (2.28) turns to

$$m\ddot{r} = m\omega^2 r = m\omega_0^2 \frac{a^4}{r^3} \Rightarrow \frac{d}{dt} \dot{r}^2 = -\frac{d}{dt} \frac{\omega_0^2 a^4}{r^2} \Rightarrow \dot{r}^2 = -\frac{\omega_0^2 a^4}{r^2} + \text{const}$$

At  $t = 0$   $\dot{r} = 0$  so  $\text{const} = \omega_0^2 a^2$  and

$$\dot{r} = \omega_0^2 a^2 \frac{r^2 - a^2}{r^2} \Rightarrow 2\omega_0^2 a^2 dt = \frac{dr^2}{\sqrt{r^2 - a^2}} \Rightarrow r(t) = a \sqrt{1 + \omega_0^2 t^2}$$

and the Coriolis force at time  $t$  is

$$2m(\vec{\omega} \times \dot{\vec{r}}) = 2m\omega \dot{r} \hat{\phi} = 2m\omega_0 \frac{a^2}{r^2} \dot{r} \hat{\phi} = \frac{4m\omega_0^3 a t}{(1 + \omega_0^2 t^2)^{3/2}}$$

### Solution # 2

In the inertial frame the point mass moves with constant velocity  $v = \omega_0 a$ . Suppose at  $t = 0$  the point was at  $x = a, y = 0$ , then at later time  $t$  it will be at the coordinates  $x = a, y = \omega_0 a t$  in Cartesian or  $r = a \sqrt{1 + \omega_0^2 t^2}$ ,  $\phi = \arccos \frac{1}{\sqrt{1 + \omega_0^2 t^2}}$ . Since the disk moves

in such a way that the point is always moving in radial direction, the angular velocity of the disk is

$$\omega(t) = \dot{\phi}(t) = \frac{d}{dt} \arccos \frac{1}{\sqrt{1 + \omega_0^2 t^2}} = \frac{\omega_0}{1 + \omega_0^2 t^2}$$

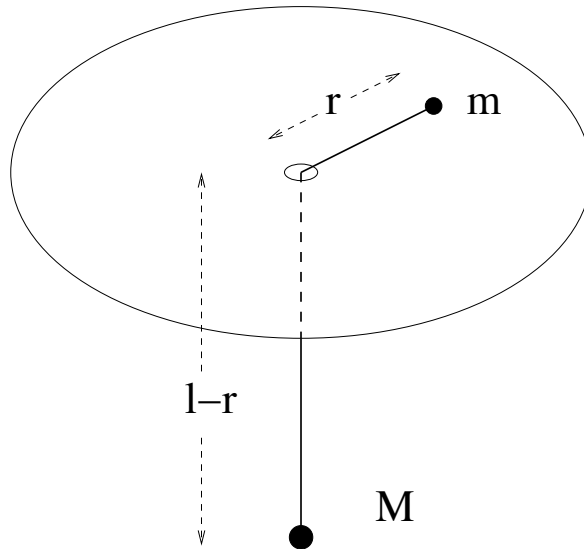
and the Coriolis force at time  $t$  is

$$2m(\vec{\omega} \times \dot{\vec{r}}) = 2m\omega\dot{r}\hat{\phi} = \frac{4m\omega_0^3 at}{(1 + \omega_0^2 t^2)^{3/2}} \hat{\phi}$$

**Problem 3.**

Two point masses  $M$  and  $m$  are connected by the (massless) rope of length  $l$ . The rope is suspended from a small hole in the frictionless table as shown below. The mass  $M$  can move (freely) only up or down while the mass  $m$  is free to move on the table surface.

1. Write the Lagrangian and Euler-Lagrange equations for this system.
2. At time  $t = 0$  the mass  $m$  is at distance  $r_0 < l$  from the hole and its velocity is  $v_0$  in the direction orthogonal to the rope. Find at which  $v_0$  the motion is circular



**Figure 1.** Projectile motion.

**Solution**

1. Let us choose  $r$  and  $\phi$  as generalized coordinates. The kinetic and potential energies are

$$T = \frac{m}{2}(\dot{r}^2 + r^2\dot{\phi}^2) + \frac{M}{2}\dot{r}^2, \quad V = Mg(r - l)$$

$\Rightarrow$

$$L = T - V = \frac{m}{2}(\dot{r}^2 + r^2\dot{\phi}^2) + \frac{M}{2}\dot{r}^2 + Mg(l - r)$$

The generalized momenta are

$$p_r = \frac{\partial L}{\partial \dot{r}} = (M + m)\dot{r}, \quad p_\phi = \frac{\partial L}{\partial \dot{\phi}} = mr^2\dot{\phi}$$

Since  $\frac{\partial L}{\partial r} = mr\dot{\phi}^2 - Mg$ ,  $\frac{\partial L}{\partial \phi} = 0$  the Euler-Lagrange equations take the form

$$\begin{aligned} \dot{p}_r &= mr\dot{\phi}^2 - Mg &\Rightarrow & (m + M)\ddot{r} = mr\dot{\phi}^2 - Mg \\ \dot{p}_\phi &= 0 \end{aligned}$$

3. From the last equation we see that  $p_\phi = mr^2\dot{\phi} = L$  is constant  $\Rightarrow \dot{\phi} = \frac{L}{mr^2}$  and the first Euler-Lagrange eqn. takes the form

$$(m + M)\ddot{r} = \frac{L^2}{mr^3} - Mg$$

Thus, in the radial direction the particle moves under the influence of the “effective force”  $F_r = \frac{L^2}{mr^3} - Mg$  which corresponds to the “effective potential”

$$V_{\text{eff}} = \frac{L^2}{2mr^2} + Mgr$$

The condition for the circular orbit is  $F_r = \frac{L^2}{mr^3} - Mg = 0 \Rightarrow v_0 = \sqrt{\frac{Mgr_0}{m}}$